



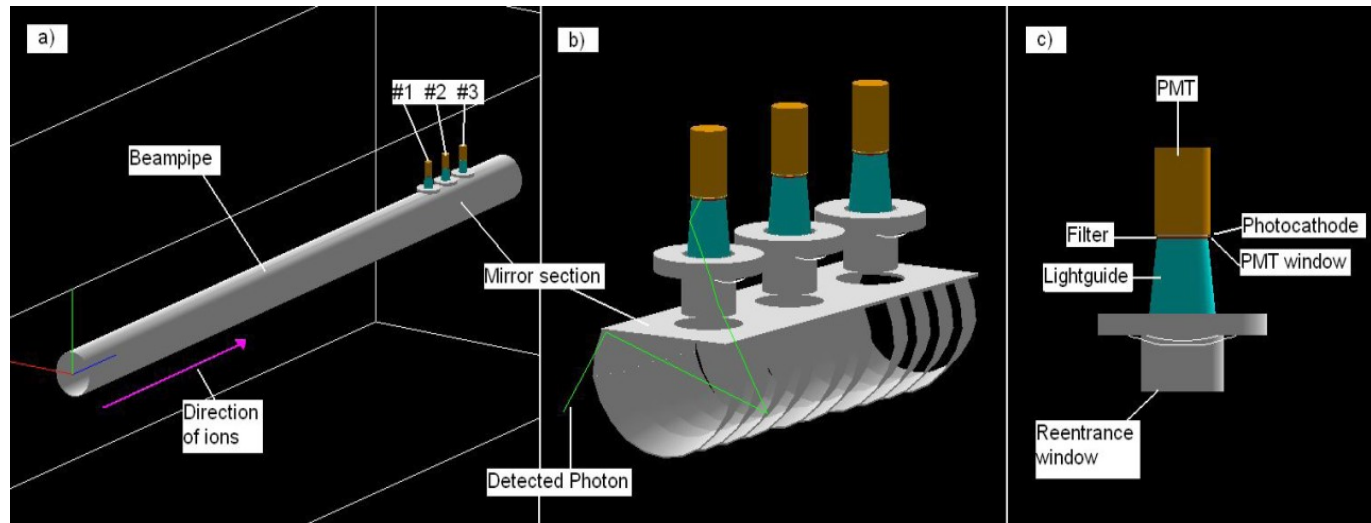
Detection systems for the $^{209}\text{Bi}^{80+}$ HFS experiment

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Mirror section

Used in the last experiment (2003)

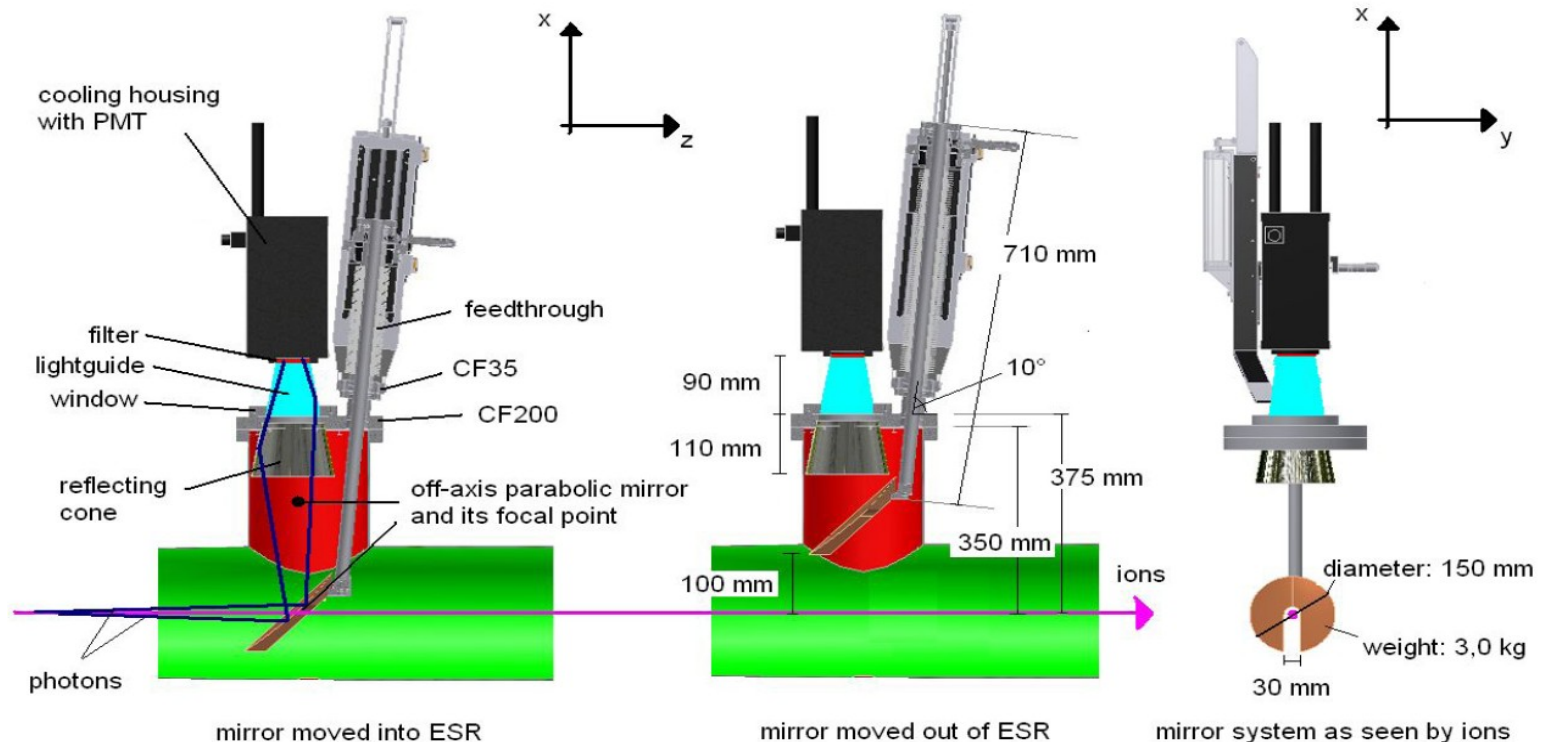


Problems because of very low overall signal rate
of max. 15 Hz on a background > 600 Hz

→ need for a dedicated detection setup, adapted to the
characteristics of the experiment

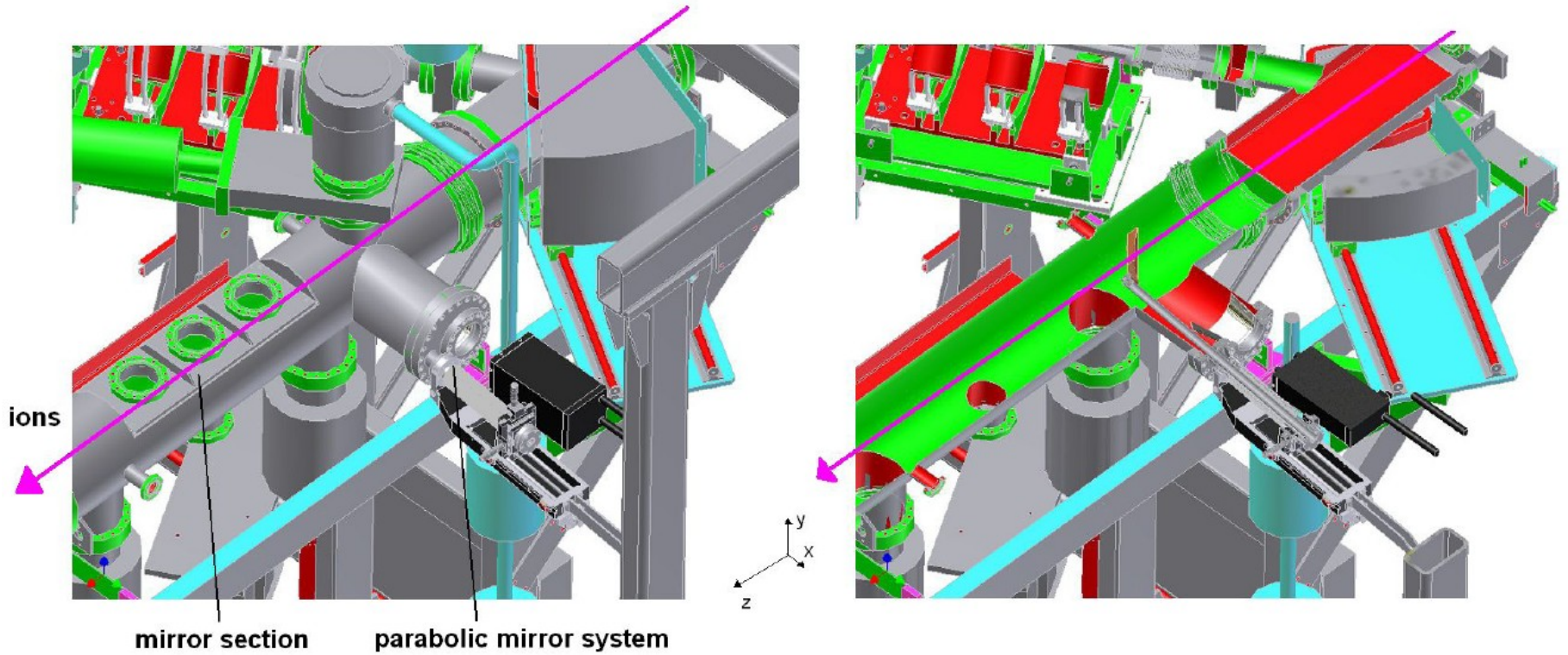
Solution: Parabolic mirror system

- Parabolic mirror with a central slit for ions, that can be moved into the beam to efficiently collect photons emitted under small angles (\approx small wavelength)

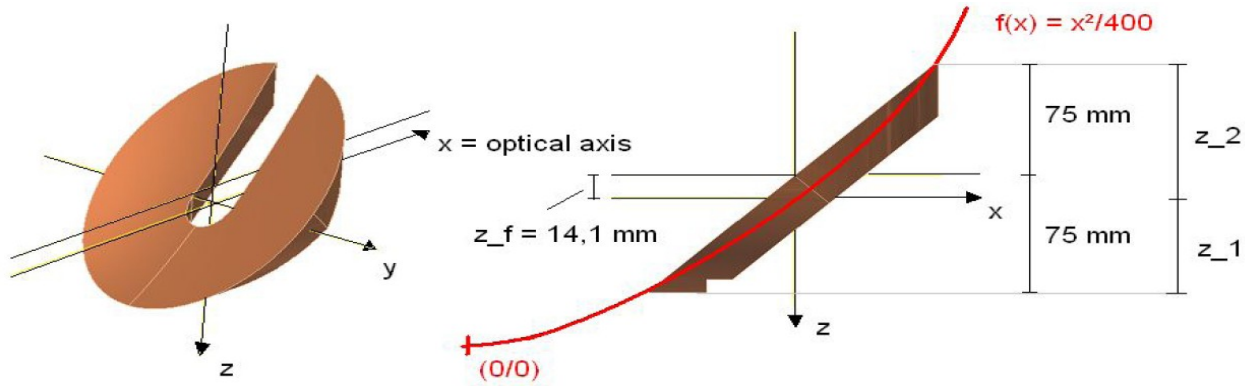




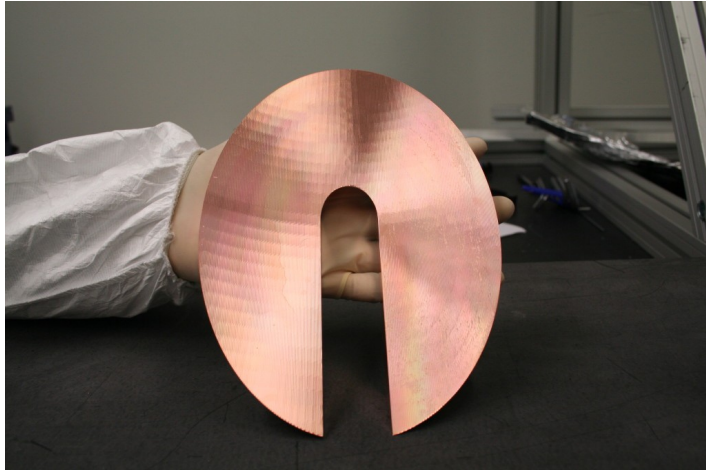
Setup at ESR



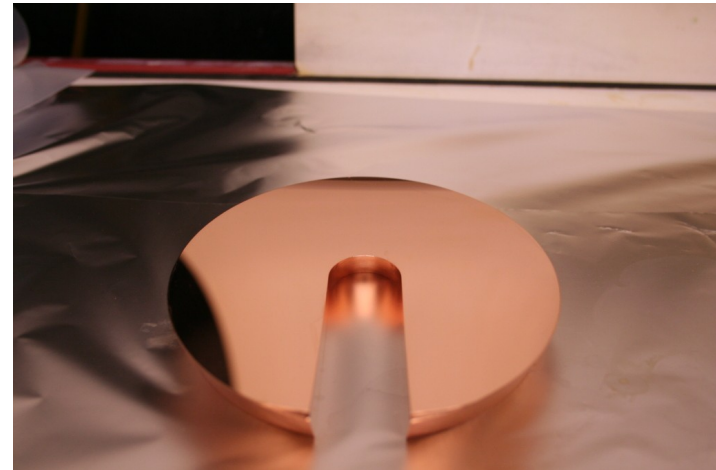
Parabolic mirror



mirror before



and after surface finish



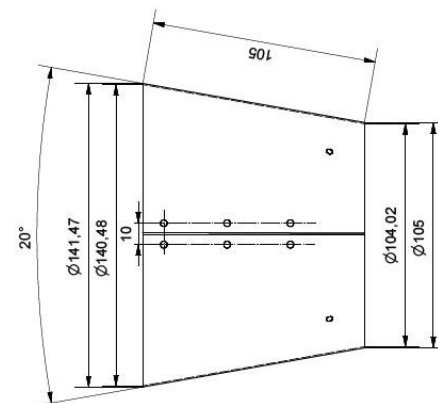
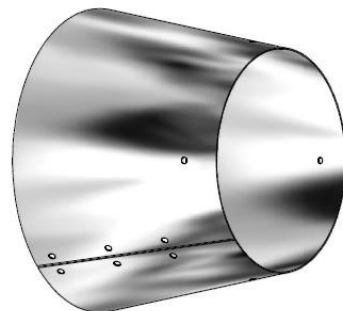
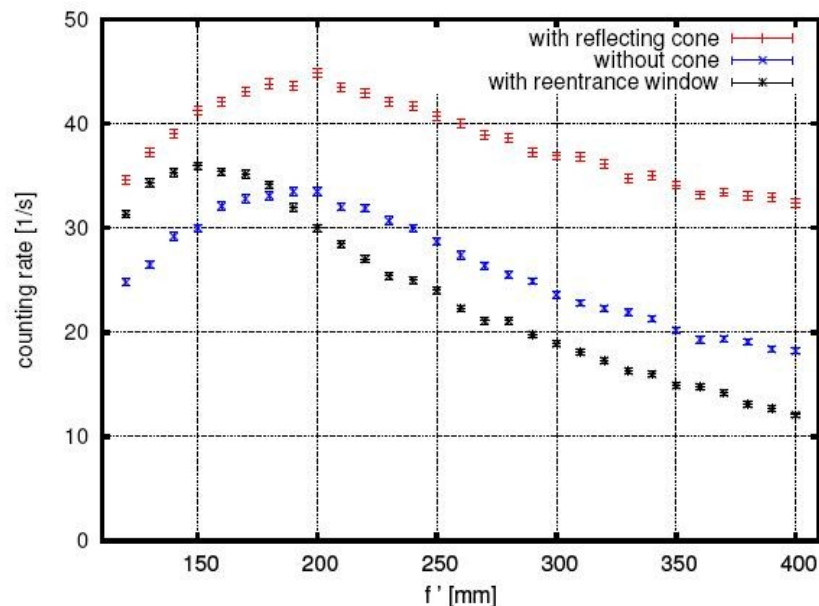
measured reflectivity for 635 nm laser light: $(97.1 \pm 1.0) \%$



Mirror characteristics



- Highest count rate at focal length of $f' = 200$ mm using an additional aluminum (Miro 2) mirror cone on the vacuum side of the exit window with $R = (89.4 \pm 0.9)\%$
- Simulated signal rates: 45 - 83 Hz, depending on exact QE of PMT and beam pipe reflectivity (background 375 - 600 Hz)

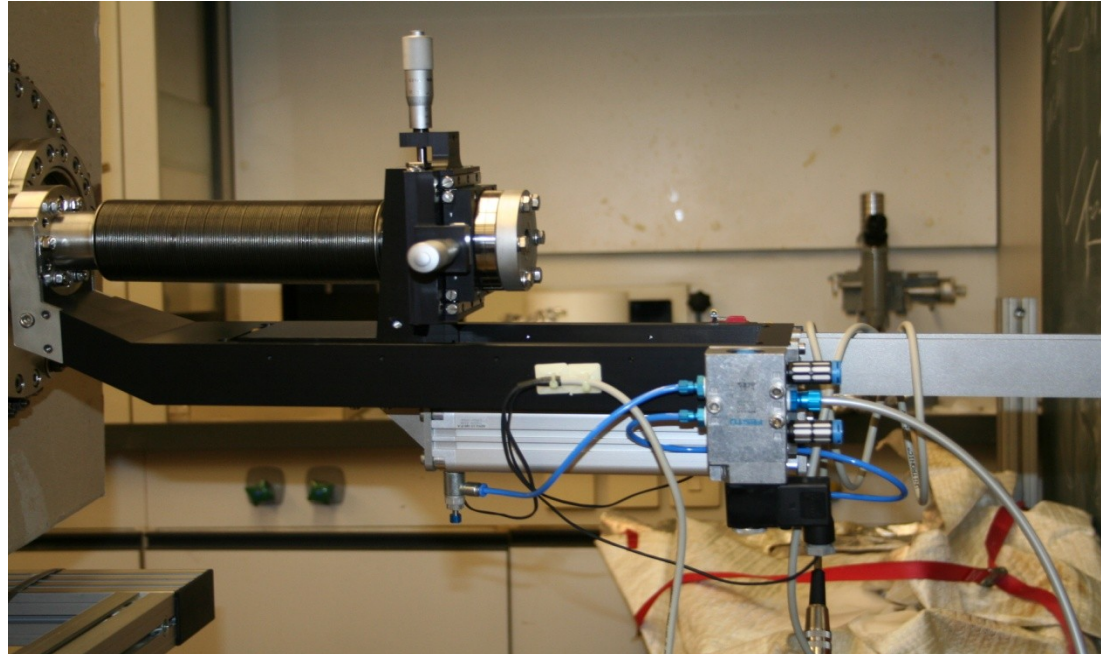




Linear feedthrough



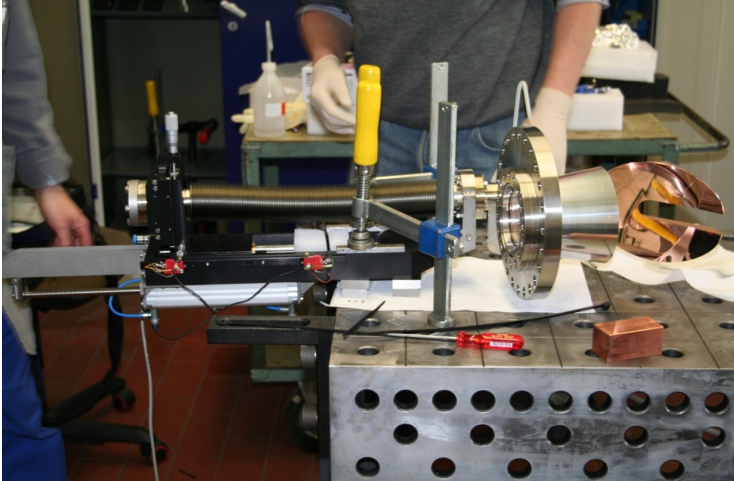
- Pneumatic system with 180 mm range to move parabolic mirror in and out



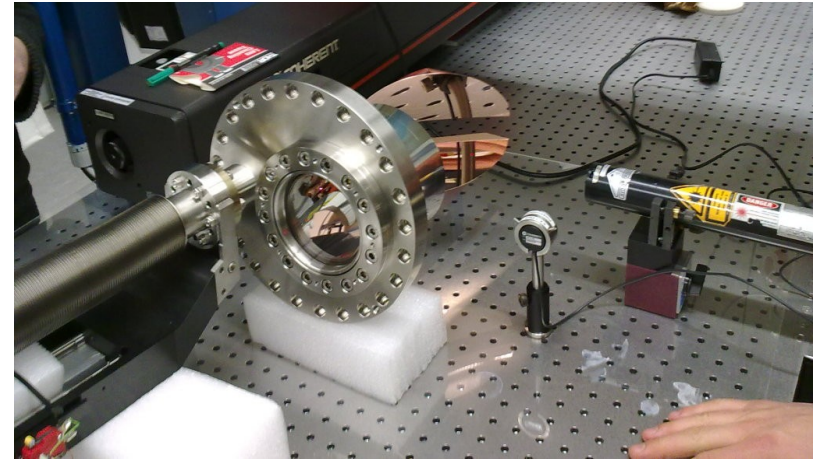
- Assembly and preliminary leak test performed in Münster



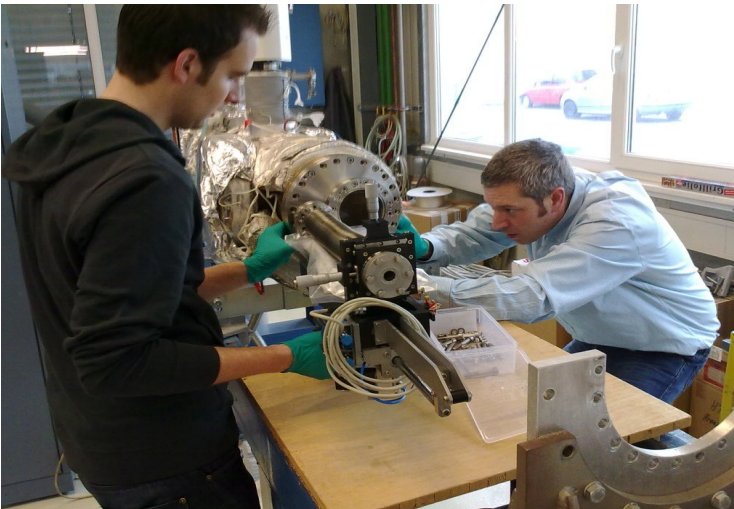
Mounting of parabolic mirror



Assembly and geometrical
adjustments in Münster



measurement of reflectivity
at GSI

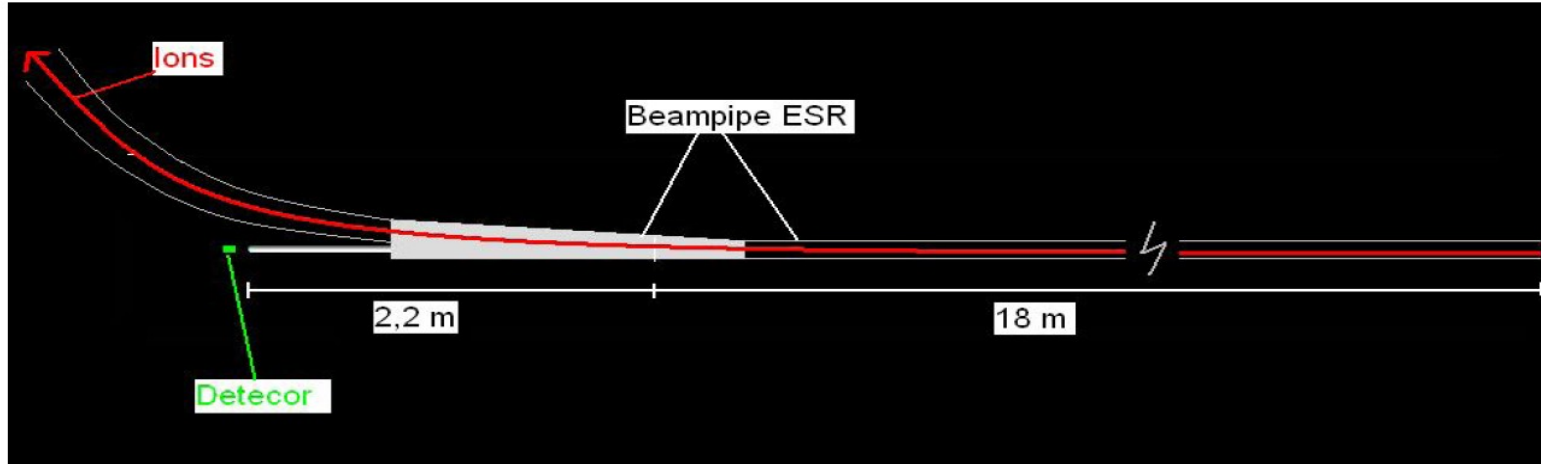


Preparation for UHV
tests at GSI

... by now the mirror has already
been mounted at the ESR



Detector in forward direction



- An additional detector can be placed at the laser entrance window in forward direction
- However, due to the small solid angle a very low count rate is expected: $S = (2.1 \pm 0.3) \text{ Hz}$ on a $\sim 30 \text{ Hz}$ background
- A suitable PMT (low noise, high QE) at this position might still be a valuable addition for diagnosis purposes
- If the HFS signal is found, it can be confirmed by this detector in a $\approx 100 \text{ s}$ measurement interval



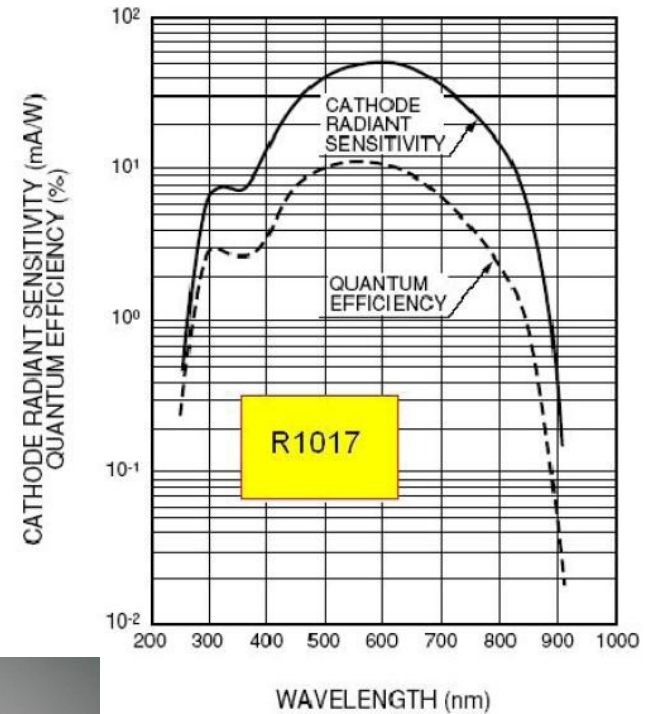
PMT types



- Requirements forward direction:
 - High quantum efficiency to be able to detect very low photon rates
 - Very low dark count rate
- Requirement Parabolic mirror:
 - Large area to detect photons reflected by the mirror
 - and of course also a high QE
- Compare PMT measurements to CPM 1993, which has a QE of 4% at 635 nm, radius 7 mm, gain $\sim 10^7$ and a dark count rate < 10 Hz (at room temperature)

PMT-R1017

- Used with parabolic mirror
- Large area: Radius = 23 mm
- Dark counts @ -20°C = 80 Hz-100 Hz
- Q.E.(theor.) @ 630nm > 10%
- Q.E.(meas.) = $2.1 \times \text{Q.E.}(\text{CPM})$
- Need to set threshold very close to the noise level due to low gain

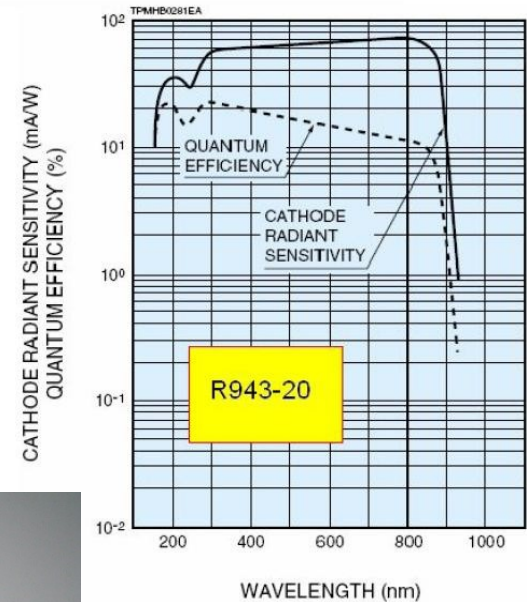




PMT R943-02



- Used in forward direction
- 10x10 mm² active area
- Dark counts @ -20°C = 15 - 20 Hz
- Q.E.(theor.) @630 nm = 14%
- Q.E.(meas.) = 2.6 × Q.E.(CPM)





Conclusions



- Parabolic mirror setup with PMT R1017 should achieve signal rates between 45 - 83 Hz on 375 - 600 Hz background
- Forward detector acts as additional diagnosis tool
- To reach the desired sensitivity, the PMT needs to be operated very near the noise level
→ although it has a smaller area, the CPM 1972 type with $QE = 8\%$ @ 630 nm could still be an interesting candidate due to its high gain of 10^7



Outlook



- Parabolic mirror:
Design of the PMT adaptor and the light guide has to be done
- Parasitic test measurements with the complete setup at GSI
- Forward direction:
Optic system with lens, light guide, filter and adaptor has to be designed
- Check if CPM 1972 could be an alternative to the PMT R1017

